

PREMIER - Instrument development of the millimetre-wave limb sounder MWLS

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History and Context

PREMIER/MWLS is designed to measure weak trace gases in the UTLS region of the atmosphere in the presence of clouds, providing a higher vertical resolution than previous instruments. Figure 2 shows the two observation bands and a select number of **observable species**.

Science objectives include dynamical stratosphere troposphere exchange processes, tropical tropopause clouds and humidity, radiative forcing, biomass burning events, industrial pollution, ozone chemistry, active nitrogen and halogen gases and general tracer measurements.

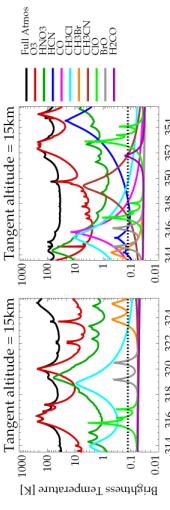


Figure 2: Simulation of the atmospheric spectrum to be observed by the MWLS (black line) and contributions from individual species (coloured lines). The dashed line denotes the detection limit due to instrument noise.

Science Rationale

The PREMIER/MWLS is designed to measure weak trace gases in the UTLS region of the atmosphere in the presence of clouds, providing a higher vertical resolution than previous instruments. Figure 2 shows the two observation bands and a select number of **observable species**.

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As part of PREMIER, the MWLS will contribute trace gas measurements in the presence of ubiquitous cirrus clouds which are opaque at infra-red frequencies, thus greatly extending the PREMIER observation range to lower altitudes. Many original STEAM-R specifications are still present in the updated instrument concept, namely

- Use of the **upper tropospheric transparency window** at 300–360GHz
- Limb-sounding geometry for high vertical resolution
- Large antenna for a narrow field of view in the UTLS region
- Fixed, full-range antenna array for rapid observations and reliability
- Dual polarisation setup allows tighter stacking of antenna beams
- Along-track sampling enables tomographic 2-D retrievals

Hardware Development

The heterodyne technique traditionally used in millimetre-wave receivers implies that a principal sideband and an image sideband are both mapped onto a single observation band. The resulting spectral confusion is either tolerated, or the image sideband is terminated by a well defined thermal load, which in return increases the noise figure of the receiver. For MWLS a novel sideband separating receiver is being developed. By exploiting the combined signal from two phase-shifted heterodyne receivers, both sidebands are returned individually, without spectral confusion.

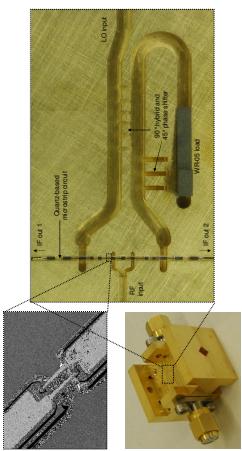


Figure 3: First prototype of the sub-harmonic image rejection mixer (SHIRM) developed for the MWLS. Top left: Anti-parallel pair of planar Schottky diodes. Right: Micro-machined cavity. Bottom left: Assembled SHIRM block.

Radiometric Performance

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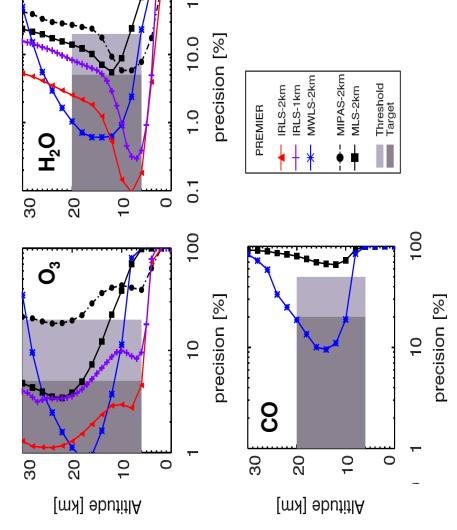


Figure 4: Measurements of sideband suppression at different frequencies.

Technical Specifications

Parameter	Unit	Value
Spectral coverage	GHz	313.5-355.6
Upper sideband separation	-	Separated sidebands
Lower 12 km	-	Double sidebands
Upper 10 km	km	6-38
Vertical Coverage	km	2-24
Arctic	km	2-24
Vertical FOV	GHz	313.6-355.6
@ 352 GHz	km	2.4
@ 324 GHz	km	2.6
Vertical sampling	-	-
Lower 12 km	km	1.5
Upper 10 km	km	2
Horizontal sampling along track	km	<-50
First satellite knowledge (ACKP)	dB	<-36
FOV sideband knowledge in orbit (ACKP)	dB	<-32
Spectral resolution	MHz	23
Baseline	MHz	6 to 12
In orbit option	-	-
Knowledge of sideband ratio	-	-
Dynamic range	dB	3-300
Radiometric sensitivity (2.5 s integration)	K	-0.5
10 MHz bandwidth	K	-0.3
23 MHz bandwidth	K	-1.0
Absolute radiometric accuracy	K	-1.0

Table 1: Detailed technical specifications of the MWLS instrument.

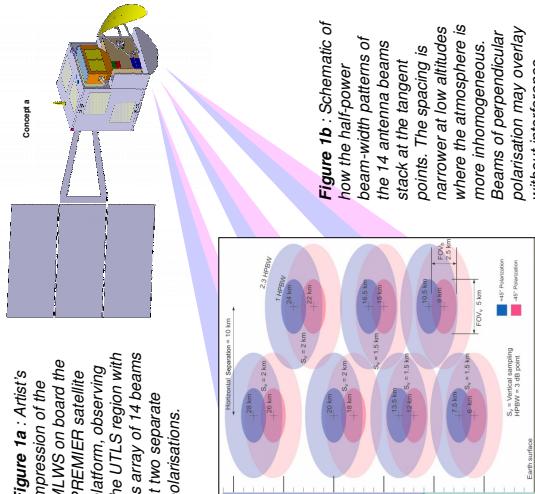
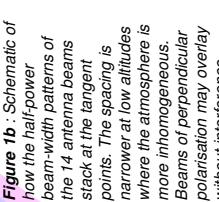
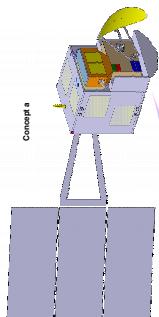


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