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Published version information

Citation: C Unsworth et al. "Characterisation of a small electrode HPGe detector." Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment, vol. 927 (2019): 293-300.

DOI: [10.1016/j.nima.2019.02.043](https://doi.org/10.1016/j.nima.2019.02.043)

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Accepted Manuscript

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PII: S0168-9002(19)30231-1
DOI: <https://doi.org/10.1016/j.nima.2019.02.043>
Reference: NIMA 61916

To appear in: *Nuclear Inst. and Methods in Physics Research, A*

Received date: 25 November 2018
Revised date: 15 February 2019
Accepted date: 16 February 2019

Please cite this article as: C. Unsworth, A.J. Boston, H.C. Boston et al., Characterisation of a small electrode HPGe detector, *Nuclear Inst. and Methods in Physics Research, A* (2019), <https://doi.org/10.1016/j.nima.2019.02.043>

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Characterisation of a Small Electrode HPGe Detector

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Abstract

Small electrode HPGe detectors in an inverted coaxial geometry are increasingly in use in applications where both high efficiency and excellent energy resolution are required. The unusual electric field configuration of these detectors results in extremely long charge collection times compared to planar and coaxial devices. In this work we have characterised such a detector using gamma-ray coincidence measurements and optimised an electric field simulation to reproduce the positional variation of detector response. We show that, alongside accurate crystal geometry and applied electric potential, a temperature correction is crucial to correctly determining appropriate charge carrier mobility parameters. This work will help to guide the future development of HPGe detectors for applications including radioactive waste assay, radio-isotope dating, and fundamental nuclear physics.

Key words: HPGe, SAGE Well, Charge collection, Charge carrier mobility

Figure 1: Schematic of the SAGe Well HPGe crystal used in this work sliced along the x axis (red). The y and z axes are shown in green and blue respectively. The 25 mm diameter p+ electrode surface is shown on the back face of the crystal in red and the passivated region surrounding it in blue, the n+ electrode covers the remaining surface of the crystal including the inside of the well. The front face of the crystal is tapered in towards the well in order to eliminate regions of very low electric field, which would otherwise lead to significant charge trapping.

1. Introduction

High-resolution gamma-ray spectrometry is used across a range of applications. These include a number of fields where both high efficiency and excellent energy resolution are critical to performance. In applications such as environmental measurements, where sample sizes are often limited, a well geometry is often used to maximise efficiency. The Small Angle Germanium (SAGe) Well [1, 2, 3] is a p-type high-purity germanium (HPGe) detector manufactured by Mirion Technologies Canberra and designed to have excellent energy resolution and very high efficiency for gamma-rays emitted by samples placed within the well.

The induced charge signals due to gamma-ray interactions at a range of positions within a SAGe Well detector have been investigated in order to characterise the charge collection behaviour. Electric field simulations and experimental measurements have been used for the characterisation with the latter being used to validate and optimise the former. The optimised field simulations can be used to predict the characteristics of other similar detectors and aid in the design of future devices.

In this work we have used a SAGe Well with a diameter of 85 mm and a length of 66 mm (Fig. 1). The crystal has a 33 mm diameter well bored into the front face to a depth of 41 mm and a 26 mm taper which reduces the diameter of the crystal to 15 mm at the front. These features help to reduce regions of extremely low field, which would otherwise lead to significant charge trapping. A small p+ electrode of 25 mm diameter is separated by a passivated region from the n+ electrode covering the rest of the crystal surface. The operating bias of -4700 V was applied to the p+ electrode. Relative to other detector geometries of similar volume, the small size of the p+ electrode provides reduced capacitance and hence electronic noise. This reduced noise helps the detector achieve outstanding energy resolution with FWHM of 0.73 keV at 122 keV and

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1.69 keV at 1332 keV. The detector was mechanically cooled by a CP5-plus cryo
 30 cooler which allows the device to be operated in any orientation, a fact which
 facilitated our characterisation measurements.

Points in the detector are described by a Cartesian coordinate system with
 the xy plane coincident with the back face of the detector and the z axis running
 through the centre of the crystal towards the front face. The origin is in the
 35 centre of the p+ electrode and the x , y , z axes run parallel to the $\langle 100 \rangle$,
 $\langle 010 \rangle$, and $\langle 001 \rangle$ crystal axes respectively (See *Fig. 1*).

2. Simulation

Signal formation in the detector was simulated using the AGATA Data Li-
 brary (ADL) [4] which was adapted for this work to incorporate the SAGe Well
 40 geometry. The simulation uses a finite difference method to solve the electric
 and weighting fields in the detector before tracking holes and electrons through
 the field using the mobility parameterisation described in [5]. The charge tra-
 jectories are then used to calculate the signal induced on an electrode using
 the Shockley-Ramo theorem [24]. The detector was initially modelled accord-
 45 ing to the nominal geometric and material specification and using the electron
 and hole mobility parameters described in [8]. Optimisations of the simulation
 parameters to match the experimental signals are described in *Sec. 6*.

Fig. 2 shows a slice of the calculated electric potential in cylindrical polar
 coordinates (rz with r in the xy plane), the electron and hole trajectories for
 50 each of the example signals we will be considering are also shown (See *Sec. 5*
 and *Fig. 4* for the equivalent experimental signals). Due to rotational symmetry
 of the detector this potential is the same regardless of the angle in the xy plane
 at which the slice is taken. The unusual field distribution in the volume sur-
 rounding the well causes the electrons (dashed blue lines) to follow the potential
 gradient towards the closest part of the n+ electrode while the holes approach
 the potential “valley” partway ($r \approx 28$ mm) between the outer detector wall
 and inside of the well. The holes then drift through this valley along the $\langle 001 \rangle$

Position	x (mm)	y (mm)	z (mm)
R1	-9.5	0.5	13.5
R2	-14.5	0.5	13.5
R3 / D1	-27.5	0.5	13.5
R4	-36.5	0.5	13.5
D2	-27.5	0.5	30.5
D3	-27.5	0.5	46.5
D4	-27.5	0.5	62.0

Table 1: Coordinates in the detector frame of positions R1 - R4 and D1 - D4

crystal axis with the trajectories from each position converging onto a single path as they approach the p+ electrode. Holes produced by interactions anywhere in the region surrounding the well, which represents the majority of the detector volume, follow a similar path as they approach the p+ electrode.

Fig. 3 shows the weighting potential for the p+ electrode together with the same charge trajectories shown on the electric potential. The weighting potential is close to zero throughout most of the detector volume until it begins to rise quickly in the vicinity of the p+ electrode ($z < 20$ mm). Together these potentials result in a range of signal shapes depending on the position of interaction in the detector. For illustrative purposes we will consider signals produced due to interactions in positions R1-4 and D1-4, the coordinates for each of these interaction positions is shown in *Table 1*. Note that points R3 and D1 are the same.

The combination of converging hole trajectories and a weighting potential concentrated close to the p+ electrode gives rise to the important features of the signal shapes from positions D1 to D4 (see *Fig. 4*). Electrons are collected quickly over a short distance while holes have a long drift with very little induced signal, then a significant induced signal as the holes approach the electrode which has a fixed shape regardless of initial interaction position.

This behaviour results in charge drift times increasing with distance from the

electrode in both the radial and z directions. Longer collection times are seen when the charge has to drift further along the common path in the z direction, or further in the radial direction before reaching the common path.

Considering the signals from the set of interaction positions R1-R4 (*Fig. 4*), on a line through the detector radius at a fixed z of 3.5 mm. Signals from positions R1 and R2, closest to the p+ electrode, have very fast (<250 ns) almost linear rising edges. Referring again to the simulated electric potential, *Fig. 2*, we see that this region of the detector has straight electric field lines with a roughly linear change in potential between the p+ electrode and the bottom of the well. Electron and hole drift distances are comparable and the charge trajectories for R1 and R2 differ only in that R2 approaches the p+ electrode at a slightly greater angle, resulting in a slightly greater rise time. Holes still make the dominant contribution to the induced signal, due to the weighting potential for the p+ electrode changing more quickly close to the electrode, but electrons do play a significant role here unlike elsewhere in the detector.

As the radius of the interaction position increases to positions R3 and R4 we see a return to behaviour seen in the well walls with electrons playing little part in the induced signal and the holes converging onto the same common trajectory before approaching the electrode.

3. Experimental Methodology

The University of Liverpool detector characterisation system, *Fig. 5*, consists of a 1 Ci ^{137}Cs source mounted inside a lead and tungsten collimation assembly. The tungsten collimator is 160 mm long with an outer diameter of 10 mm and a 1 mm diameter hole, this sits inside an array of lead blocks. The position uncertainty produced by the collimator in this measurement varied from 1.2 mm diameter at the front of the detector to 2.2 mm at the back. The entire assembly is mounted on top of a pair of linear stepper motors allowing it to be moved in two dimensions to a precision of 0.1 mm.

This scan table is combined with a fully digital data acquisition system

Figure 2: An r -slice of the electric potential (Volts) calculated by ADL for the SAGe Well detector, due to rotational symmetry of the detector this is the same for a slice taken at any angle in the xy plane. Electron (dashed blue) and hole (red) trajectories are shown for the seven example positions discussed in the text, the simulated interaction positions are marked with black circles.

Figure 3: An r - z slice of the weighting potential calculated by ADL for the SAGe Well detector p+ electrode, due to rotational symmetry of the detector this is the same for a slice taken at any angle in the xy plane. Electron (white) and hole (red) trajectories are shown for the seven example interaction positions discussed in the text, the simulated interaction positions are marked with black circles.

Figure 4: The top panel shows simulated signals for selected positions in the detector using the nominal detector parameters and previously published values for hole and electron mobility, see *Sec. 2*. The bottom panel shows the experimental mean signals measured during the coincidence scan for the same positions, see *Sec. 3*. See *Fig. 2* and *Fig. 7* for the corresponding interaction locations.

Figure 5: A schematic of the University of Liverpool Detector Scanning Table mounted with a SAGe Well detector in coincidence mode. The source, primary collimator and lead collar are moved by the stepper motors while the rest remains stationary. The red line indicates an example path of a valid coincidence Compton scatter.

using Caen V1724 100 MHz 14 bit digitisers to capture detector charge signals. This system allows the positional response of a detector to be investigated by interrogating it with the beam of 662 keV gamma rays[9]. The rate of gamma rays coming from the collimator is ≈ 1000 per second.

The scanning system can be operated in two modes, *singles* and *coincidence*. In singles scanning mode the collimator is raster scanned across the whole of the detector with the only spatial information coming from the position of the collimator when an interaction occurs. This allows identification of the xy position of the *first* interaction undergone by a gamma ray. This method does not allow location of the interaction in the z direction nor does it provide any constraint upon the total number of interactions undergone before the full energy is deposited. Despite these limitations the singles scan data are crucial to establishing the shape, position, and orientation of the detector crystal in the frame of the scanning system.

The coincidence scan method uses in addition a secondary array of collimating lead blocks with a thickness of 80 mm and 1.5 mm gaps created by

plastic spacers. The gaps are aligned with BGO scintillation detectors which identify gamma rays that Compton scatter through 90° at defined z positions and subsequently interact in one of the BGO detectors. The geometry of the secondary collimators and detectors resulted in a z position uncertainty from 2 mm at large radii of interaction to 3 mm near the centre of the detector. The triggering electronics were configured to read out all events in which interactions occurred in both SAGe Well and BGO detectors within a coincidence window of $2 \mu s$ to cover the observed range of rise times in the detector.

Combining the information from the secondary detector with the collimator position allows the full three dimensional localisation of *single-site* interactions. If a number of signals are collected from each position it is possible to form a mean signal and hence deconvolve the underlying detector response from the random electronic noise. The rate of such coincident interactions varies with position but is typically less than one event per minute compared with a random coincidence trigger rate of the order of 100 events per minute. This necessitates the use of offline event selection techniques to identify the events of interest.

Conservation of energy and momentum ensures that for a fixed gamma-ray energy a scatter through 90° will deposit a fixed energy in each of the primary and secondary detectors. In the case of 662 keV gamma rays the values are 374 keV in the HPGe and 288 keV in the BGO. *Fig. 6* shows a plot of the BGO energy versus the HPGe energy for events measured in time coincidence. The events of interest can then be selected and the background reduced with gates on both energies as shown in the figure. The width of the gates applied depends on both the collimator geometry and the energy resolution of the detectors, in this measurement our gates were 374 ± 12 keV in the HPGe and 288 ± 40 keV in the BGO. Prominent lines in the background of *Fig. 6* represent the 662 keV photopeak and 511 keV annihilation photons in the germanium with random background events in the BGO. Further background suppression can be achieved by limiting the Ge-BGO time difference to a range of values consistent with being due to the scattering of a single gamma ray. See *Sec. 5* for further discussion of the Ge and BGO time difference.

Figure 6: Energy deposited in HPGe and in BGO during the SAGe Well scan. The feature marked by the red box represents Compton scattering through 90° leaving 374 keV in the germanium and 285 keV in one of the BGO detectors. Prominent vertical lines represent the 662 keV photopeak and 511 keV annihilation photons in the germanium with random background events in the BGO.

Figure 7: 662 keV photopeak counts as a function of position for the front (left) and side (right) scans. The coordinates have been translated into the detector frame where the origin is in the centre of the p+ electrode and the x , y , and z axes run parallel to the $\langle 100 \rangle$, $\langle 010 \rangle$, and $\langle 001 \rangle$ crystal axes respectively. Example points from the coincidence scan, discussed in Sec. 5 are indicated.

Following the application of initial time and energy gates the signals are interpolated linearly between the measured points at 10 ns intervals down to 2 ns samples. They are then shifted in time to align the point where they reach 10% of their maximum to a fixed sample number and normalised to equal height before an initial mean signal for this position is formed. The final stage of filtering is to compare each individual signal to this initial mean and measure the RMS difference between the two, signals with large differences are rejected and a final mean signal is formed from those that remain.

4. Singles Scan Results

Fig. 7 shows the positional variation of 662 ± 2 keV photopeak counts observed when the SAGe Well detector was scanned with the collimated ^{137}Cs

165 source. The left image shows a scan from the front of the detector and the right shows a scan from the side, in both cases the collimator was held at each position for 4 seconds. The most probable way for a 662 keV photon to leave its full energy in germanium is by first Compton scattering and then leaving its remaining energy at another location through photoelectric absorption. This
170 results in a reduction in photopeak counts at the detector edges in both scans as the probability of the photon scattering out of the sensitive volume before depositing its full energy increases.

The well in the centre of the detector is visible in both scans as a reduction in intensity because there is less sensitive germanium material present in the path
175 of the gamma-ray beam. Other regions of reduced intensity are the result of gamma rays interacting with attenuating material before they reach the detector. See for example the thin ring in the centre of the front scan, a consequence of scattering in the walls of the cryostat endcap, and bands of reduced intensity in the side scan, caused by the material used to physically support the crystal.

180 Also indicated on *Fig. 7* are the interaction locations for the example signals that will be discussed in *Sec. 5*. Points R1 to R4 lie on a radial line relatively close to the p+ electrode at $z = 13.5\text{mm}$. Points D1 to D4 lie at a fixed radial position on a line through the depth of the detector and parallel to the z axis.

5. Coincidence Scan Results

185 In the coincidence measurement the first parameter to be studied was the time difference between the signals measured in the SAGe and BGO detectors. The left side of *Fig. 8* shows the time difference between triggers generated in the both detectors during the coincidence scan for all x and y at each z positions, the black line shows all events generating a trigger and the coloured lines show
190 the events selected for mean signal formation at each value of z . On the right side time difference for two BGOs triggered by coincident 511 keV gamma rays from a ^{22}Na source is shown, this distribution has a FWHM of 23 ± 2 ns.

The BGO detectors generated consistent signal shapes which were not de-

Figure 8: Histograms of the SAGe-BGO time difference obtained during the coincidence scan (left). The black line represents all events generating a coincidence trigger, the coloured lines represent signals passing all gates and being used to form mean signals. The BGO-BGO timing response to 511 keV annihilation photons from a ^{22}Na source is provided for comparison (right), the FWHM of this distribution was 23ns.

pendent on the position of interaction. We therefore assume any variation in the
 195 SAGE-BGO trigger time difference greater than the BGO resolution to be due
 to variation in the SAGE response. This distribution reveals the large range of
 charge collection times observed in the SAGE detector, with total collection time
 increasing with increasing distance from the p+ electrode up to a maximum of
 1.6 μ s for events near to the detector front face.

200 The signals in the SAGE detector will now be investigated for a range of
 different positions within the detector. For each mean signal formed *Fig. 9*
 shows the rise time for the initial (*left*) and final (*centre*) parts of the pulse
 (2% to 30% and 30% to 98% of its height respectively). Also shown is the
 mean Ge-BGO trigger time difference for events contributing to a mean signal
 205 (*right*), as a function of the position of interaction. The x axis shows the radius
 of the interaction position and the colour and shape of the markers indicate the
 z position.

The rise time of the initial part of the rising edge shows little variation
 through most of the detector volume as shown by the clustering of events around
 210 400 ns. However there is a strong dependence on radius for events occurring at z
 positions closer to the p+ electrode, with shorter rise times seen at smaller radii.
 The final part of the rising edge again shows little variation through most of
 the detector volume but has a slight dependence on z for events close to the p+
 electrode. The SAGE-BGO time differences on the other hand reveal a strong
 215 dependence on z of the time before a trigger signal is generated by the SAGE
 detector.

These data support the simulated charge collection behaviour described in
Sec. 2, holes produced by a gamma-ray interaction will drift a long way through
 the detector before inducing any significant signal on the collecting electrode.

220 The mean signals produced by the method described in *Sec. 3* are aligned
 relative to the SAGE detector trigger time, in order to reveal the true variation in
 charge collection time it is necessary to shift them according to the mean SAGE
 BGO time differences of the contributing signals. This method produces mean
 signals with the correct shape and timing relative to the BGO signal as shown

Figure 9: Shape parameters for signals induced in SAGE detector as a function of position. Time is measured signals to go from 2% to 30% of their height (left), from 30% to 98% of their height (centre), and the mean Ge-BGO trigger time difference (right) for events contributing to the mean signal at each position. The x axis indicates the radial position of each mean signal, the z position is indicated by the colour and marker type.

225 in *Fig. 4* for the example signals R1 to R4 and D1 to D4.

6. Investigating Parameters of the Simulation

Fig. 4 shows the experimental mean signals discussed in *Sec. 5* together with simulated signals for the same positions. The simulated signals show qualitatively the same behaviour, with fast rising signal for interactions close to the p+ electrode and increasing charge drift time as distance from the electrode increases.

The total charge collection times for the simulated signals are however much shorter than seen in experiment. For example, charge collection for an interaction at point D4 located at $z = 6$ mm takes $1.1 \mu\text{s}$ compared with $1.6 \mu\text{s}$ in experiment. In particular the initial part of charge collection, where charge carriers are far from the p+ electrode and the induced signal is close to zero, is significantly faster in the simulation.

Broadly, the rate of charge collection is determined by the electric field in the detector and the mobility of charge carriers being collected. The electric field at each point in the detector is determined by the detector crystal geometry, the potential at the electrode, and the internal electric field due to stationary space charge. Details of the geometry were checked against the intensity profiles obtained through singles scanning (see *Fig. 7*) and while this measurement revealed slightly smaller dimensions than the specification the discrepancy can be explained by the existence of surface dead layers which are not visible in the scan.

Impurity concentrations were quoted by the manufacturer to within 10% at the front and back of the crystal, a linear gradient along the z axis was assumed between these points and no radial variation was included in the model. A linear impurity gradient along the z axis results in a uniform field in the bulk of the detector where the field due to space charge dominates the overall electric field. The linear variation of charge collection time as a function of z , seen in *Figs 8 and 9 (right)* supports the hypothesis of an impurity gradient which is

approximately linear in z . We were not able to determine if there is any variation
 255 in impurity concentration with xy , but any such variation would be expected to
 contribute mainly to the field in the xy plane and would not therefore explain
 the rate of charge collection along the z direction.

In order to establish if uncertainties in the impurity measurements at each
 end of the crystal could explain the observed charge collection times we ran the
 260 simulations again while shifting the front and back impurity concentrations by
 up to 20%. *Fig. 10* shows the results of these simulations for the induced signal
 from interactions at point D4. While there is a significant change in charge
 collection time, the maximum effect of a 20% shift is only enough to increase
 the charge collection time to $1.2 \mu\text{s}$, still significantly less than the $1.6 \mu\text{s}$ seen
 265 in experiment.

We thus conclude that realistic uncertainties in the crystal impurity concen-
 tration are not enough to explain the observed discrepancy in charge collection
 times.

The hole and electron mobility parameters used for the initial simulation
 270 were obtained by a fit to experimental data from the MINIBALL array of coax-
 ial HPGe detectors, described in *Ref [8]*. These values have had success in sim-
 ulating the signal shapes generated in coaxial HPGe detectors from the AGATA
 array [10]. While alternative parametrisations of charge carrier mobility have
 produced different values [11], comparisons with data from coaxial HPGe detec-
 275 tors have shown relatively little sensitivity of simulation performance to choice
 of mobility parameter [12, 13].

Ref [8] does not give the temperature of the HPGe crystals when the study
 was performed but the liquid-nitrogen-cooled MINIBALL and AGATA crystals
 typically vary between temperatures of 95 K and 100 K under normal conditions
 280 [14]. The temperature of the SAGe crystal in this study was 113 K, a value which
 was chosen in manufacture to optimise energy resolution. Since this difference
 is relatively large, and the temperature dependence of mobility is stronger in
 weaker electric fields, a significant temperature correction will be required here.

In order to establish if the longer charge collection times seen in experiment

Figure 10. Simulated signals from point D4 using a range of values for the impurity concentration at the head (front) and tail (back) of the detector.

could be due to this temperature difference we ran a series of simulations with
adjusted mobility parameters. As the charge collection time in the SAGe Well
is dominated by the hole mobility these signals show little sensitivity to electron
mobility. As such we focussed this study on hole mobilities only. The difference
between experiment and simulation was most pronounced for interaction posi-
tions with the longest drift distances so we used exclusively point D4, near the
front face of the detector, for this comparison.

Studies on the relationship between mobility and temperature for germanium
found a dependence on $T^{-1.6}$ for electrons and $T^{-2.3}$ for holes [15, 16]. This
dependence equates to a reduction in hole mobility of 33% when temperature
increases from 95 K to 113 K.

Fig. 11 shows the simulated signals where the hole mobilities along the
 $\langle 100 \rangle$ and $\langle 111 \rangle$ axes were both changed by between +10% and -35%. The
simulated signal with -30% hole mobility provides a good match to experiment
for both the total charge collection time ($\approx 1.6 \mu\text{s}$) and the time for the signal
to reach 10% of its height ($\approx 1.2 \mu\text{s}$). It is therefore reasonable to conclude that
the crystal temperature differences are the dominant factor in the longer charge
collection times seen in this SAGe well detector. With suitable temperature
corrections there is excellent agreement between simulated signals and those
observed in this detector.

This underlines the importance of temperature corrections to mobility for
accurate simulation of signal shapes, a point which is likely of increased impor-
tance given the proliferation of inverted-coaxial type detectors with relatively
weak fields [17, 18]. Furthermore, the increasing popularity of mechanically
cooled detectors which may be expected to have a wider range of crystal tem-
peratures than liquid nitrogen cooled devices, will also increase the importance
of such corrections.

Figure 11: Simulated signals from point D4 using a range of values for the detector hole mobility in the $\langle 100 \rangle$ and $\langle 111 \rangle$ direction. The -30% shift is equivalent to that expected from a $\mu \propto T^{-2.3}$ relationship as temperature rises from 97.5 K to 113 K.

7. Acknowledgements

This research was supported by the United Kingdom Science and Technology Facilities Council (STFC UK) grant ST/N003543/1.

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